



CLARREO Pathfinder Mission Overview and its Intercalibration Capabilities

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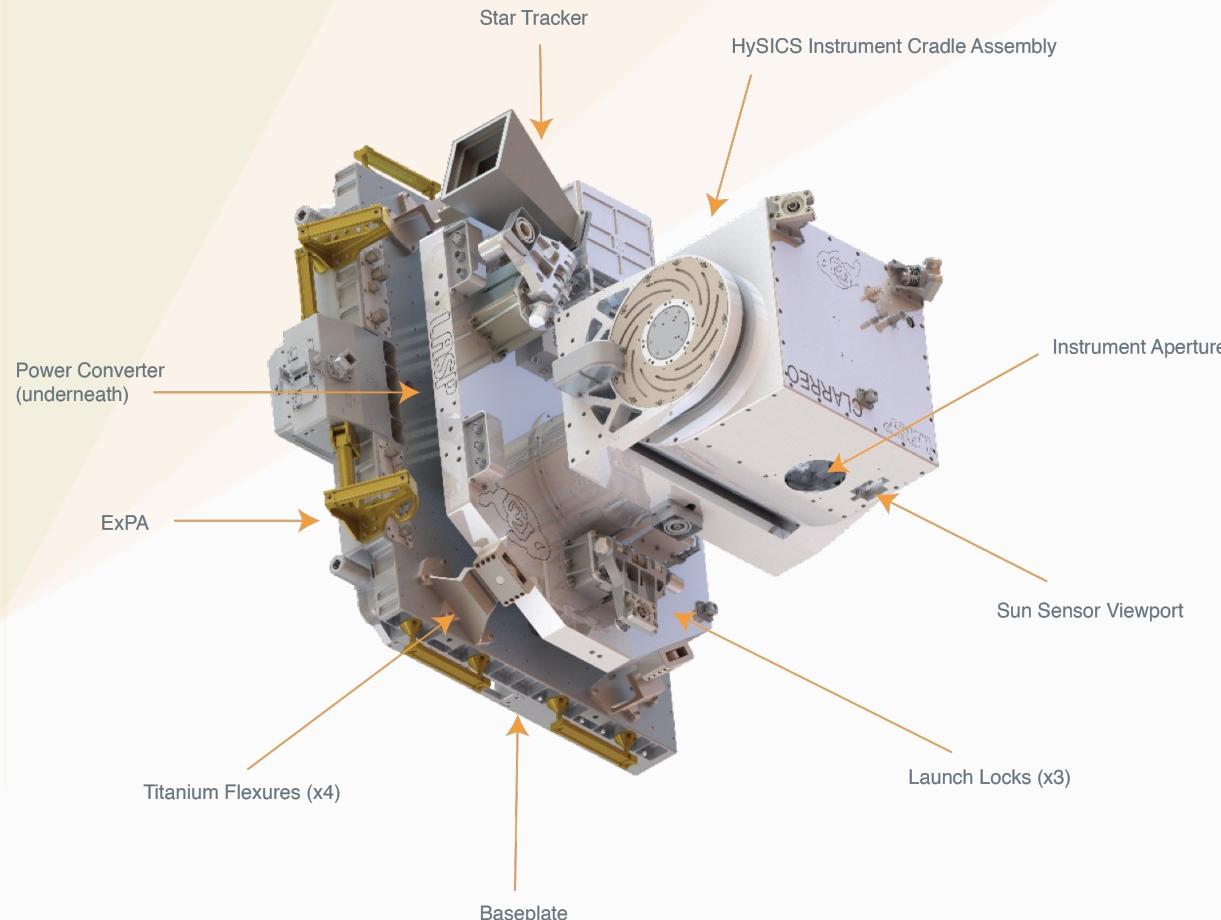
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CLARREO Pathfinder Payload



HySICS: HyperSpectral Imager for Climate Science



Push-broom spectrometer

Spectral Range	350 nm - 2300 nm
Spectral Sampling	3 nm
Radiometric Uncertainty	0.3% (1-sigma)
Swath Width	10° (70 km nadir)
Spatial Sampling	0.5 km
Platform	ISS

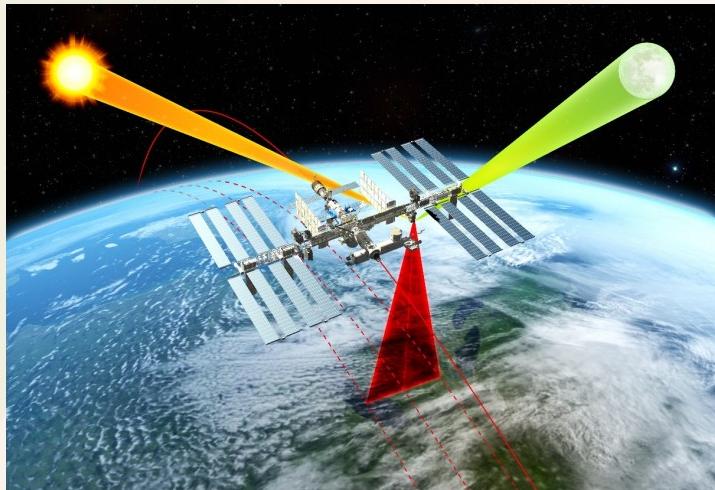
<https://clarreo-pathfinder.larc.nasa.gov/>



CPF Science Objectives

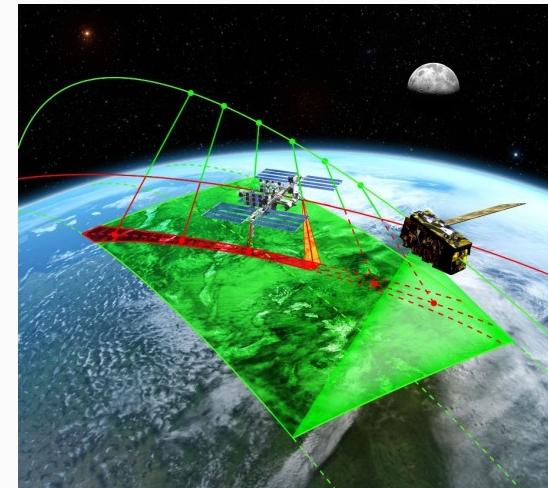


Objective #1: High Accuracy SI-Traceable Reflectance Measurements



Demonstrate on-orbit calibration ability to reduce reflectance uncertainty by a factor of **5-10 times** compared to the best operational sensors on orbit.

Objective #2: Inter-Calibration Capabilities



Demonstrate ability to transfer calibration to other key RS satellite sensors by inter-calibrating with CERES & VIIRS.

	Objective #1	Objective #2
Uncertainty	Spectrally-resolved & broadband reflectance: $\leq 0.3\%$ (1σ)	Inter-calibration methodology uncertainty: $\leq 0.3\%$ (1σ)
Data Product	Level 1A: Highest accuracy, best for inter-cal, lunar obs Level 1B: Approx. consistent spectral & spatial sampling, best for science studies using nadir spectra	Level 4: One each for CPF-VIIRS & CPF-CERES inter-cal. Merged data products including all required info for inter-cal analysis

Intercalibration: A multidimensional data matching challenge!

Spatial/Temporal Matching

- Spatial pixel resolution difference
- Non-zero temporal gap
- Mitigation: Use larger intercalibration footprints

Spectral Differences

- Scene dependent biases
- Accurate SBAF estimates needed for correction
- Hyperspectral measurements with <4nm sampling desired

Viewing/Solar Geometry

- May induce systematic biases
- Scene dependent corrections

Polarization Sensitivity

- Differing polarization sensitivity between Reference and Target
- Results in additional uncertainty in measuring polarized radiances
- Scene- and angle dependent

Miscellaneous

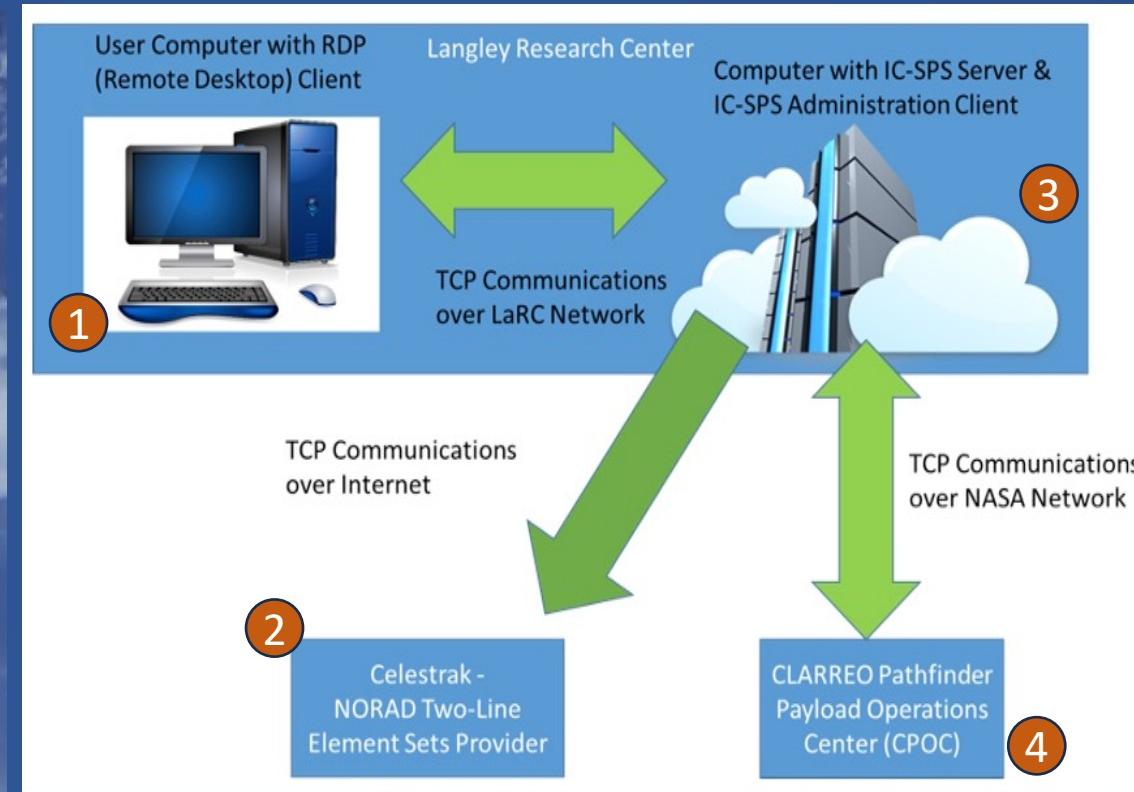
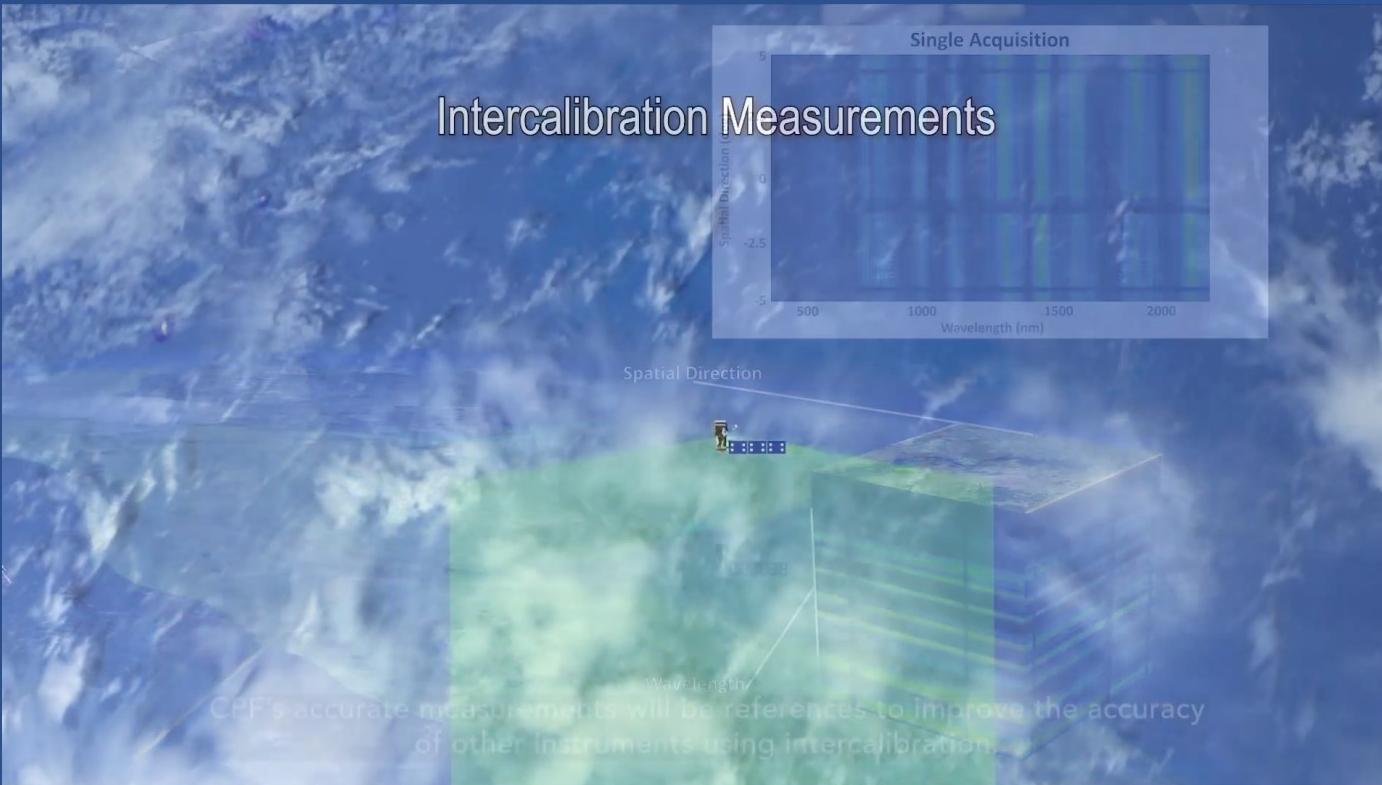
- Disparity in usage of reference solar spectra
- Scan-angle calibration dependency
- Sensor non-linearity

- In an idealized intercalibration setup, both the reference and target satellite instruments would acquire measurements that perfectly match in ***time, space, spectral response, and viewing and solar geometry***
- CPF has developed ***novel algorithms to systematically analyze data matching uncertainties*** and substantially mitigate their impacts to achieve ***unprecedented intercalibration accuracy***

CPF as a SI-traceable Intercalibration reference

- CPF exhibits the most advanced amalgamation of *spectral range*, *spectral/spatial resolution*, and *radiometric accuracy*, positioning it as the future benchmark for in-orbit intercalibration
- CPF's 3-nm hyperspectral sampling enables the creation of pseudo multispectral channels that perfectly align with the spectral response functions of the target sensor, eliminating the requirement for SBAF
- *CPF will demonstrate a state-of-the-art intercalibration methodology (0.3% uncertainty at k=1) to accurately transfer the benchmark CPF calibration reference to CERES and VIIRS instruments*

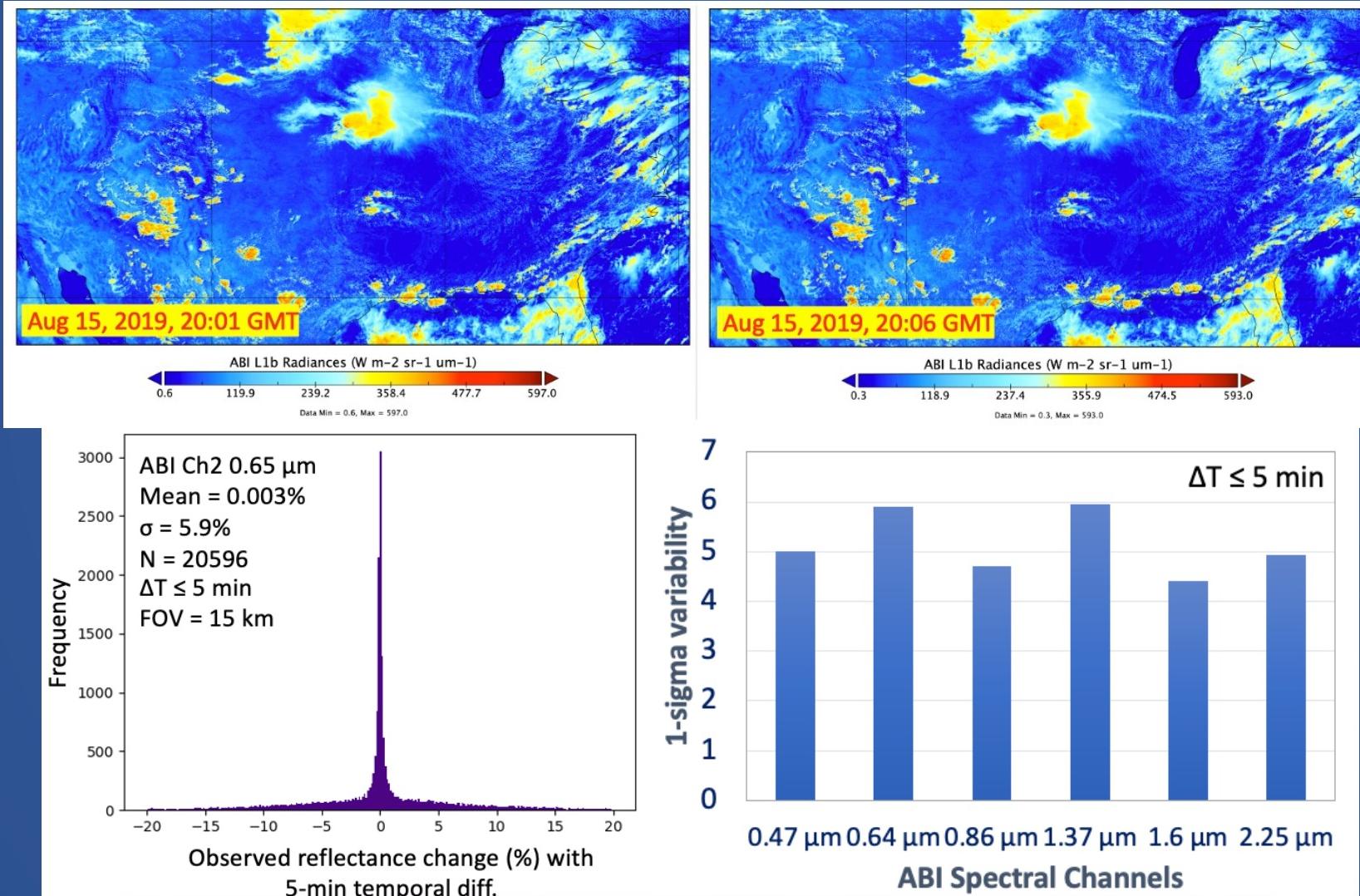
Intercalibration Science Planning System (IC-SPS)



- IC-SPS is a custom developed software tool at LaRC to accurately predict intercalibration opportunities and provide that information to CPF Operations Center
- Other variants:
 - LASICS: Generalized intercalibration event predictor- critical tool for the *Satellite Severe Convection Research Group at LaRC* (successfully identified every convective cell observed by MODIS and AMSR-E over the US from 2002-2011)
 - GCP: Ground Control Point predictor to support CPF geolocation validation
 - ARCSTONE

Spatial/Temporal Data Matching Strategy

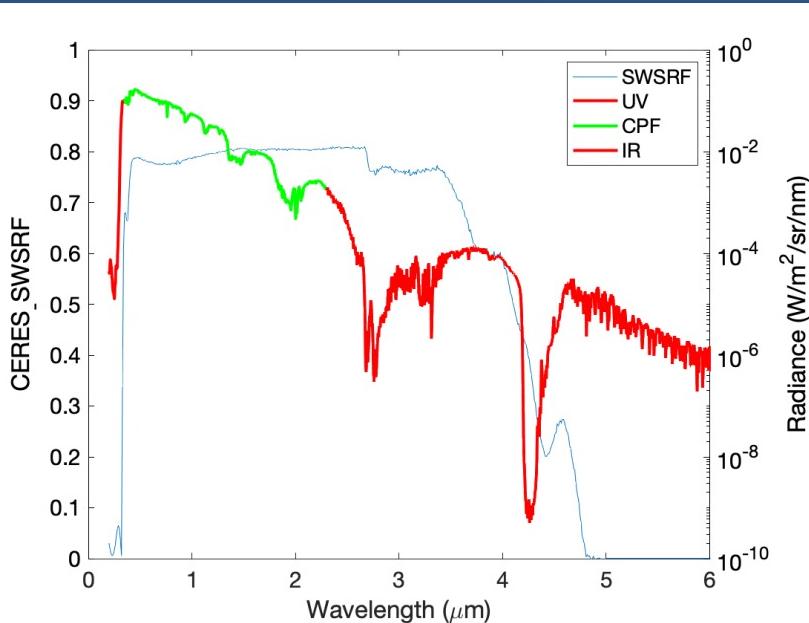
- Comprehensive and systematic evaluation of spatial and temporal matching uncertainties, considering spatial footprint size, scene homogeneity, and spectral channels as variables
- Extension of 2008 Wielicki et al. study



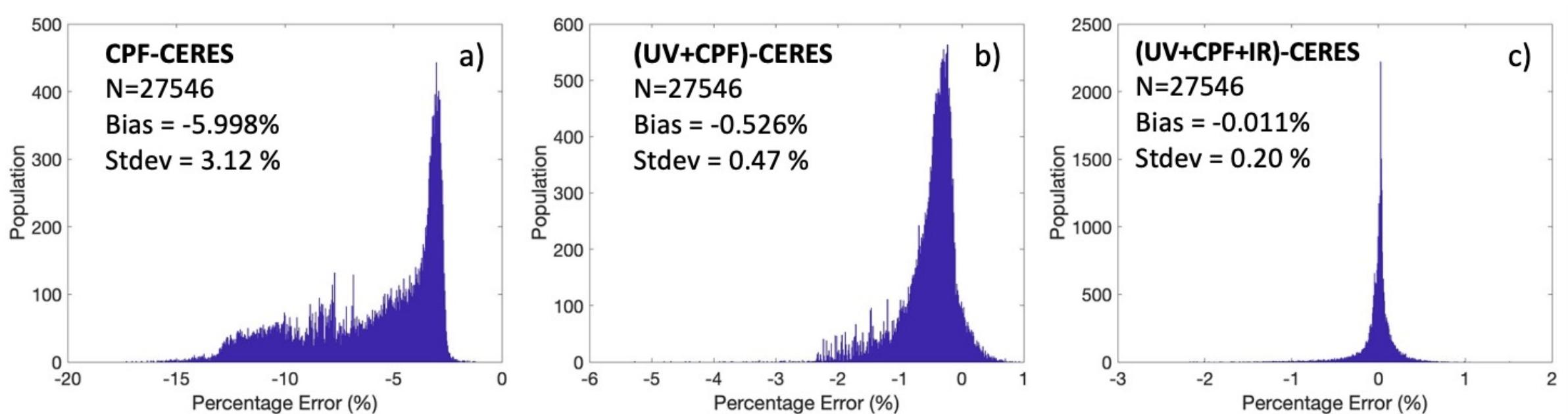
$$u = \frac{\sigma}{\sqrt{N}}$$

σ determines req. num. of intercal samples to limit data matching uncertainty below a given threshold (targeted $u = 0.1\%$)

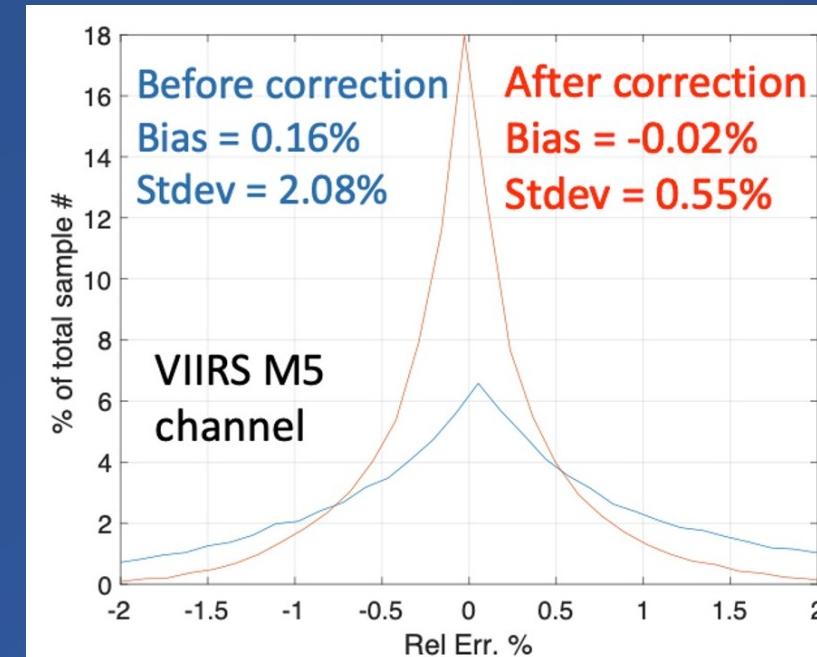
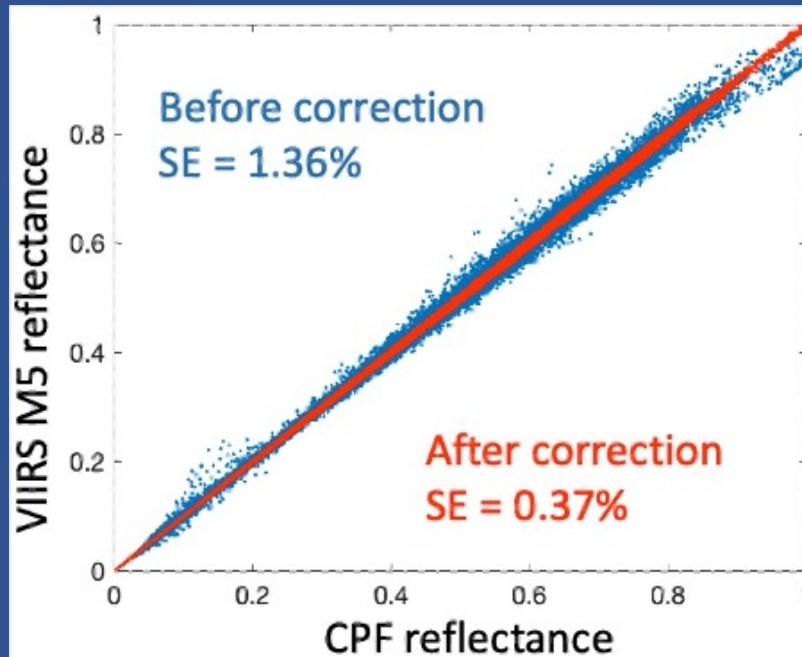
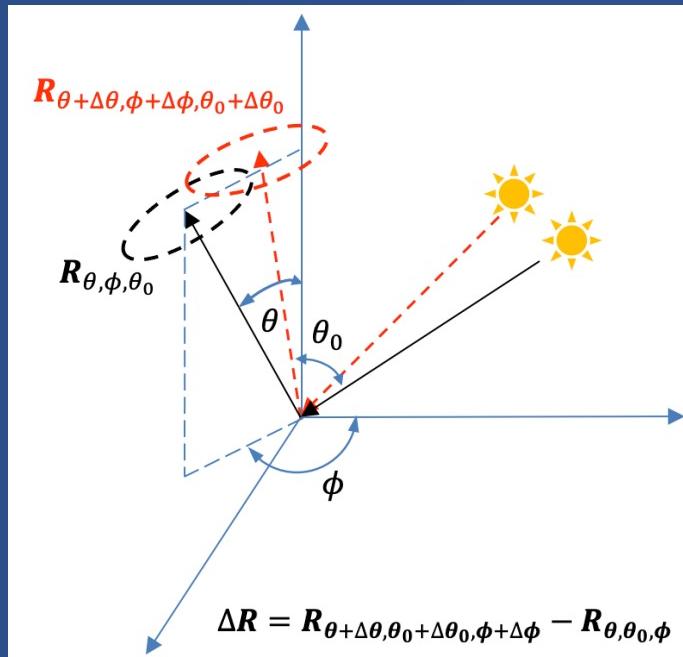
PCRTM-based Spectral Extension



- CPF measurements must be extended to 200 nm – 5 μm to intercalibrate CERES SW channel
- Leverages spectrally redundant information available in the CPF-measured portion using Principal Component Analysis and utilize pre-established *spectral correlation relationships* among wavelengths to extend the CPF spectrum below 350 nm and above 2300 nm
- Anticipated $1-\sigma$ uncertainty < 0.1%

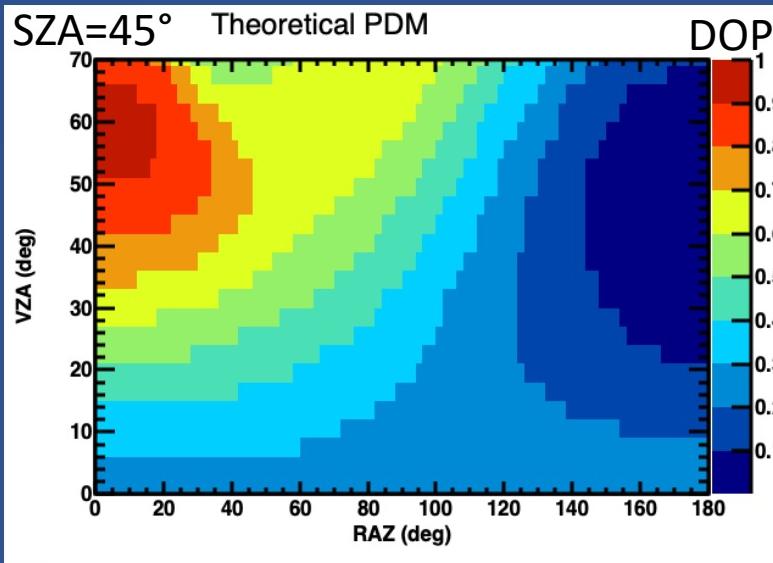
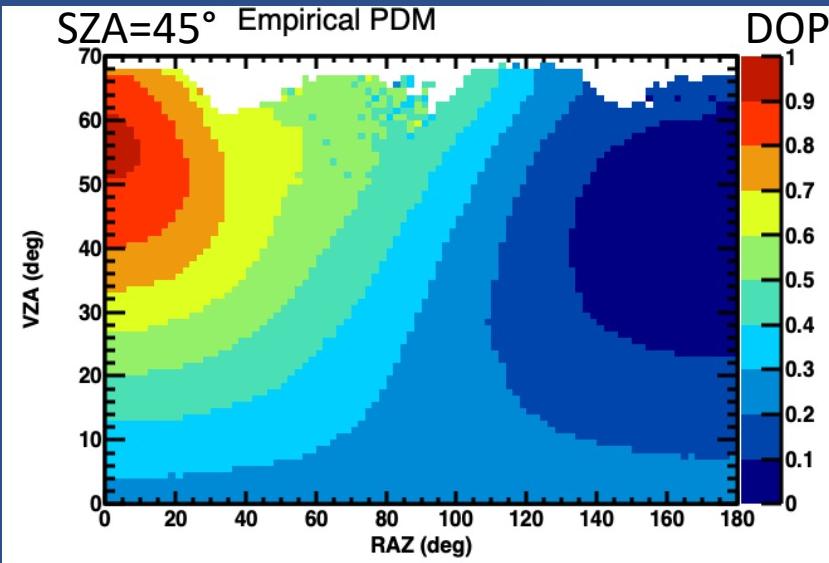


PCRTM-based Angular Correction



- CPF team has developed a PCRTM-based algorithm for angular adjustment
- Extensive training database consisting of millions of CPF-like spectral radiances encompassing diverse surface types and atmospheric conditions and SZA-VZA-RAA combinations that cover the entire range of possible viewing and solar angles
- Leverages **spectral correlation relationship** to account for differences between observations acquired at slightly different sun-view geometries
- Significant reduction of bias and noise after angular correction (uncertainty goal = 0.1%)

Novel Polarization Distribution Models (PDMs)



Disparity between diattenuation coefficient of CPF and VIIRS result in differing reflectance measurements

- Contribute to random and systematic uncertainty
- PDMs will be used for identifying low polarized intercalibration footprints

Empirical PDM for Clear-Sky Ocean (0.67 μm)

- Constructed using PARASOL/POLDER Data
- 3 wavelengths: 490, 670, and 865 nm

Theoretical PDM for Clear-Sky Ocean (0.67 μm)

- Simulated using Adding-Doubling Radiative Transfer Model
- Covers all VIIRS bands

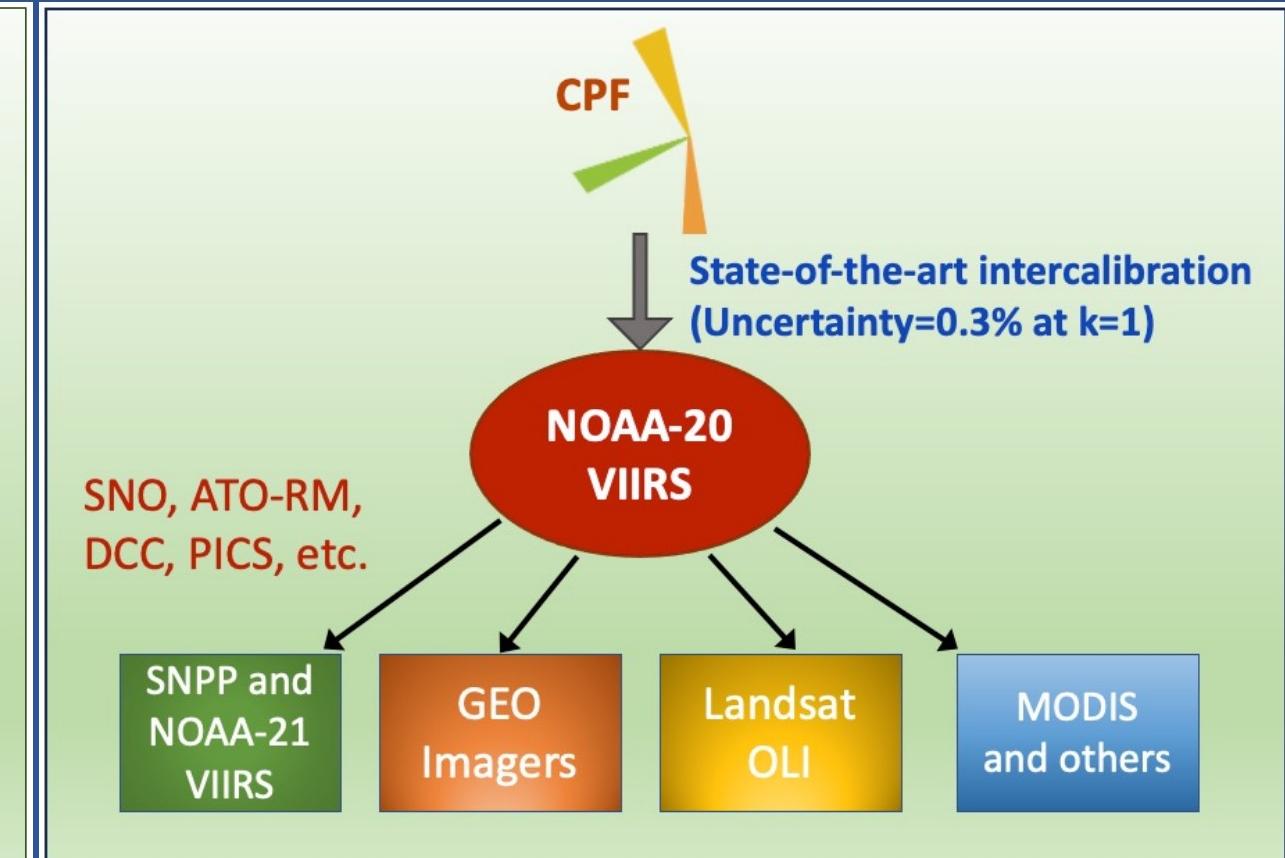
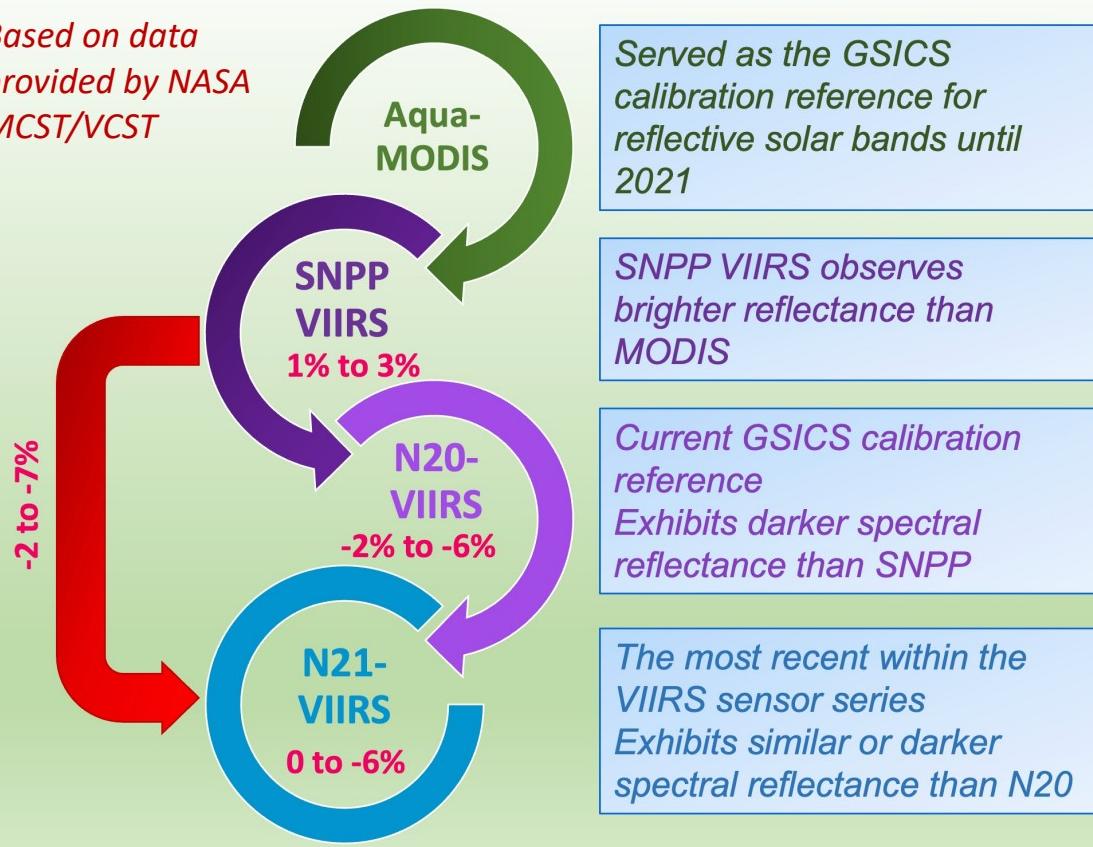
*Goldin and Lukashin (2016)
Goldin et al. (2019)
Sun et al. (2018)
Sun et al. (2015)*

PDM Application Module

- Uses VIIRS scene characterization info from L2 files, identifies correct PDM LUT and retrieves DOP/AOLP estimates from ePDMs & tPDMs
- Can be used to support several applications outside the CPF project

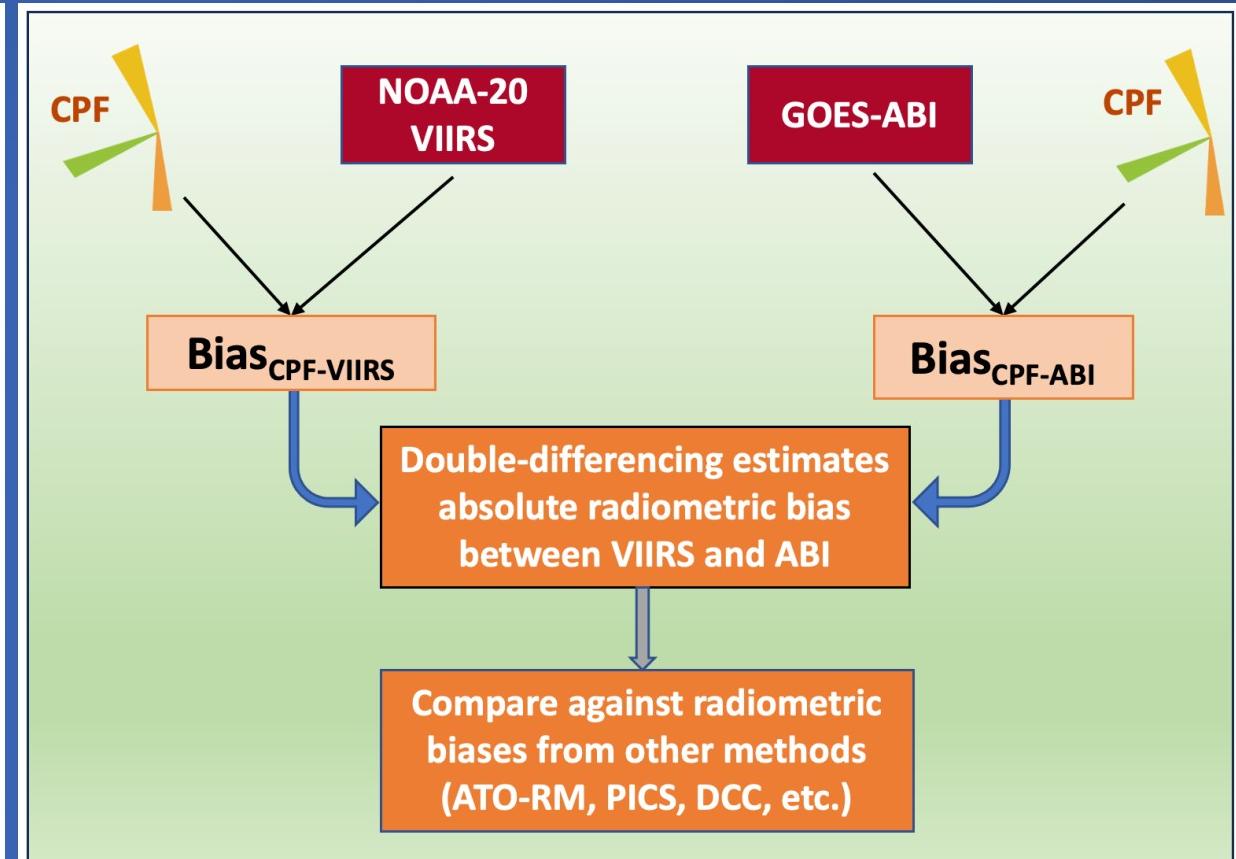
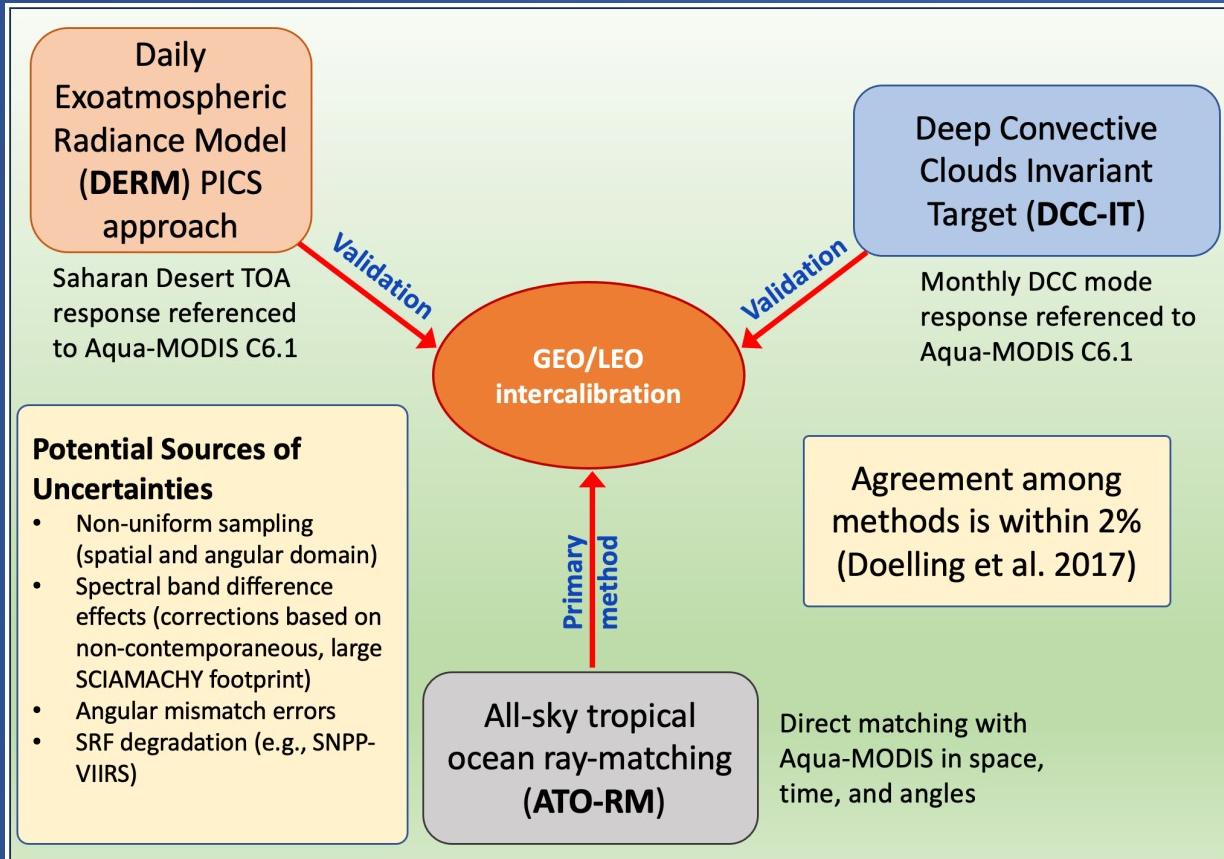
CPF will enable other Earth observing missions to surpass their original capabilities

Based on data provided by NASA MCST/VCST



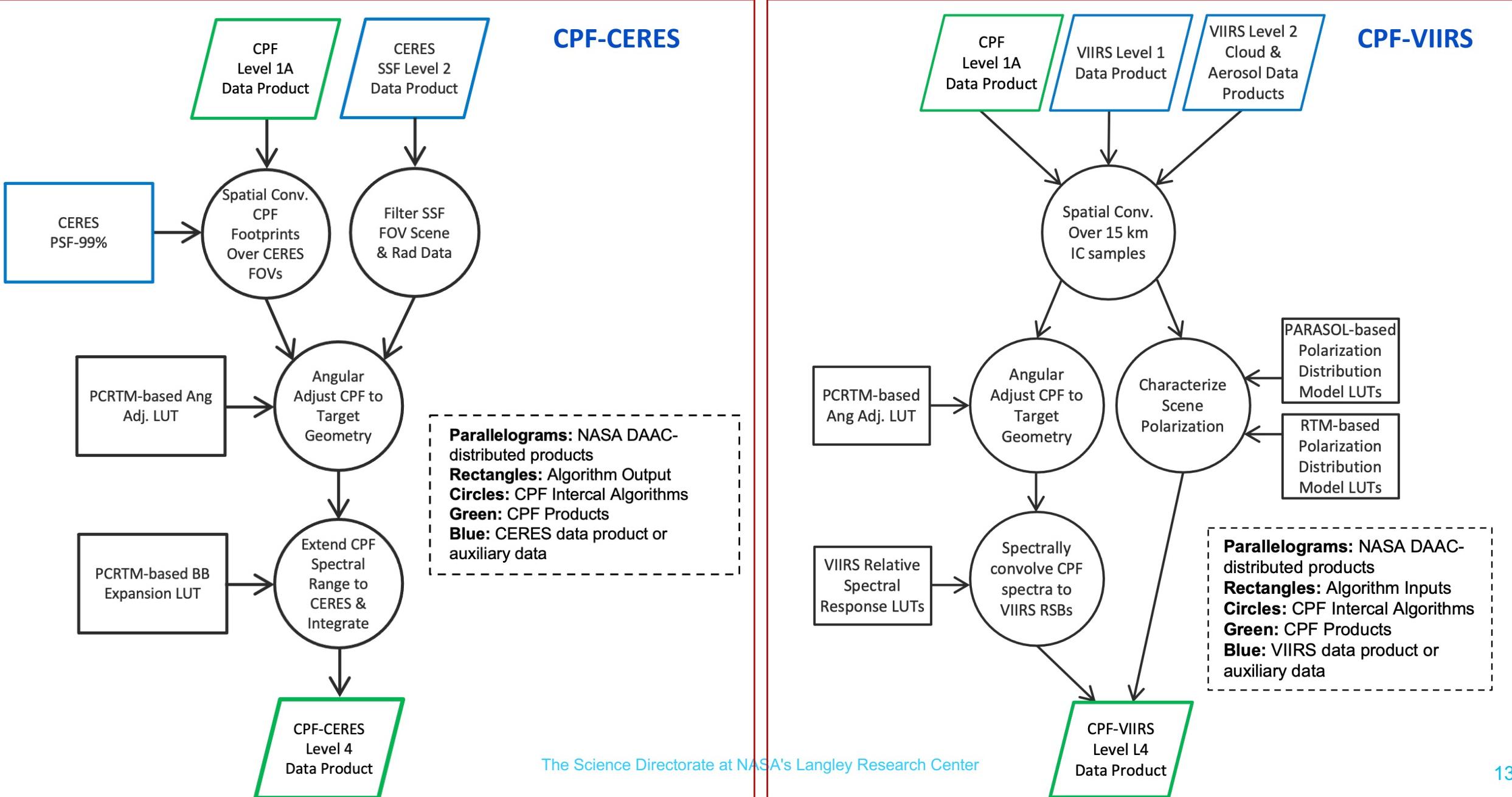
- Improved reference instrument for satellite intercalibration community
- Aid the community with an independent assessment of the radiometric accuracy of VIIRS sensors (onboard SNPP, NOAA-20, and NOAA-21 platforms)
- CPF calibration transfer to several other orbiting sensors (including MODIS) via direct approach or using VIIRS as a transfer radiometer (slightly higher uncertainty)

CPF supports validation of existing methods



- Existing intercalibration methodologies can be validated by concurrently (same month) calibrating two instruments against CPF
- CPF will help split uncertainty sources and optimize intercalibration methods for different wavelengths
- Ultimately, these methods can be used to transfer CPF reference to future as well as past instruments

Intercalibration Algorithm Flowcharts



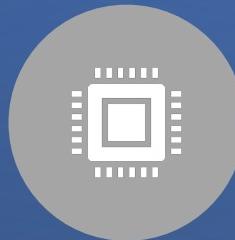
Summary and Path Forward

CPF stands out for cutting-edge fusion of spectral coverage, spectral/spatial resolution, and radiometric accuracy

CPF's innovative approaches to address the challenges of multidimensional data mismatching between CPF and target enables the attainment of intercalibration accuracy that was previously unprecedented



*Maintain collaboration with GSICS and advocate for global adoption of CPF as **preferred reference***



*Expand the utilization of CPF's innovative intercalibration methods/tools (**LASICS, PDMS**) to additional sensors and use cases*



*Enhance the spectral and radiometric characterization of Pseudo-Invariant Earth Targets (e.g., key desert sites, Deep Convective Clouds) - **Potential for additional data product***



Collaborate with CERES ICGC in maintaining Langley global leadership within intercalibration community

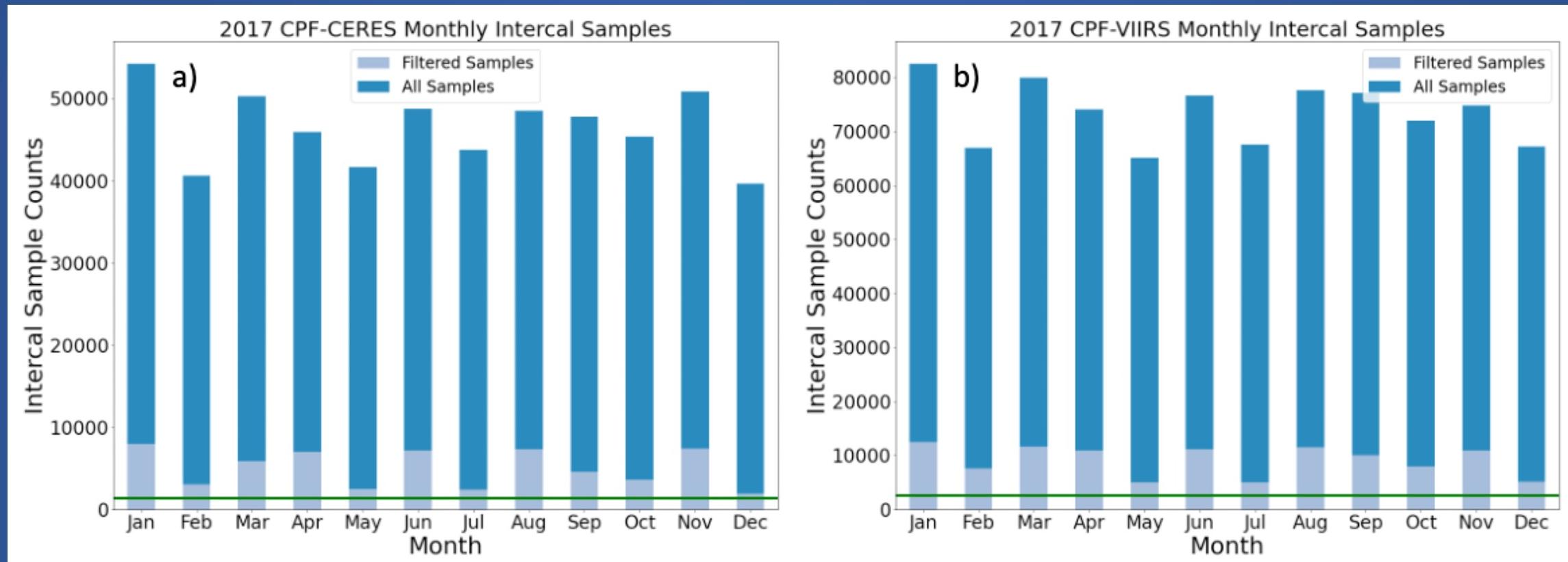


*Aid in the enhancement of the Lunar Calibration models by supplying high-accuracy Lunar reflectance measurements obtained from CPF - **Potential for additional data product***



*Anchor existing NASA Investments (e.g., MODIS) with the SI-traceable Calibration Standard of CPF, leading to an enhanced multi-decadal climate data record - **Potential for additional data product***

Intercalibration Sampling Estimates from low-fidelity simulation data for year 2017



Sample selection criteria

- a) SZA<60° and VZA <60° to ensure high signal-to-noise ratio;
- b) 5°<RAA<175° to avoid hotspot and sun glint conditions;
- c) a spatial homogeneity factor of less than 0.2 for visible wavelength (0.65 μm) to exclude extreme heterogeneous scenes;
- d) spatial field of view coverage of greater than 95%;
- e) maximum allowable time difference of 10
- f) DOP<0.1 for VIIRS

Green line represents minimum required sample size to meet uncertainty threshold